



Preparing for the next generation of (radio) observatory operations

Vanessa Moss

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IMAGE CREDIT: A. CHERNEY

Key takeaway points

- **1. People:** *Role of the technical astronomer*
- **2. Automation:** *Automate all the things*
- **3. Communication:** *Communication is the source of all the world's problems*
- **4. Overview:** *The big picture is just as important as the details*
- **5. Testing:** *There is no test like an observational test*
- **6. The shadowy place**

(O7.3) Preparing for the next generation of radio observatory operations

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Abstract. Operations of radio telescopes around the world have evolved significantly over the past decade, due to the dramatic increase in telescope complexity, available technology and data rate. In order to progress smoothly into the era of the Square Kilometre Array and other global-scale projects, it is important to bring together the lessons learned from existing observatories (such as the importance of automation for handling complex systems) and use these to shape the next generation of observatory operations. The challenges to be faced are spread across the realm of operations, such as rapid scheduling, configuration of the telescope system, visualisation of the complex raw data, data quality assessment of multi-dimensional data, "big data" management, automated processing and fast transfer to long-term archives. I will outline challenges faced and addressed in my previous role as Head of Apertif Telescope Operations, where I managed operations during the final commissioning stage of the Apertif upgrade to the Westerbork Synthesis Radio Telescope, as well as in my current role as Head of ASKAP Science Operations with CSIRO Astronomy and Space Science. I will also give examples of challenges encountered and solutions I have helped to develop for other radio telescopes including the Australia Telescope Compact Array (ATCA), the Low Frequency Array (LOFAR), and the Parkes 64m radio dish. Based on my various experiences in operations of different instruments, I will offer insight into some of the key considerations needed to develop a good model for future SKA operations.

1. Introduction

In 2032, thirteen years from now, we will celebrate 100 years of radio astronomy. The field has come a long way since Jansky's first detections of cosmic static in the 1930s, and the current generation of radio telescopes have been immensely complex to build, operate and understand. We are now at the frontier of the new generation of radio astronomy, with the Square Kilometre Array looming on the horizon, as well as existing in a time where technology is rapidly changing on unprecedented timescales. It is thus a good time to reflect on the current successes, and limitations, of the way we conduct radio observatory operations in order to help shape the path towards smooth operations of the next generation of radio telescopes. In the sections below, I will highlight various aspects of operations from my experience working with several different radio telescopes around the world. The lessons learned based on these experiences are critical to keep in mind as we move towards the SKA and beyond.

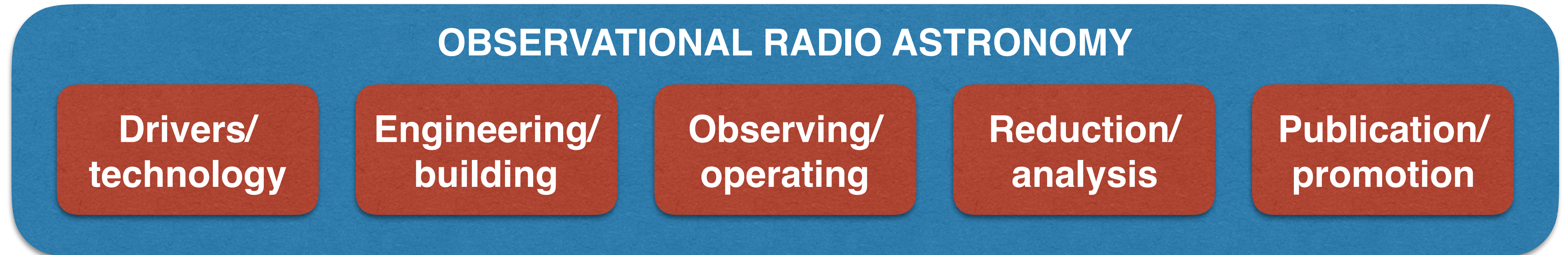
1. People

Role of the technical astronomer

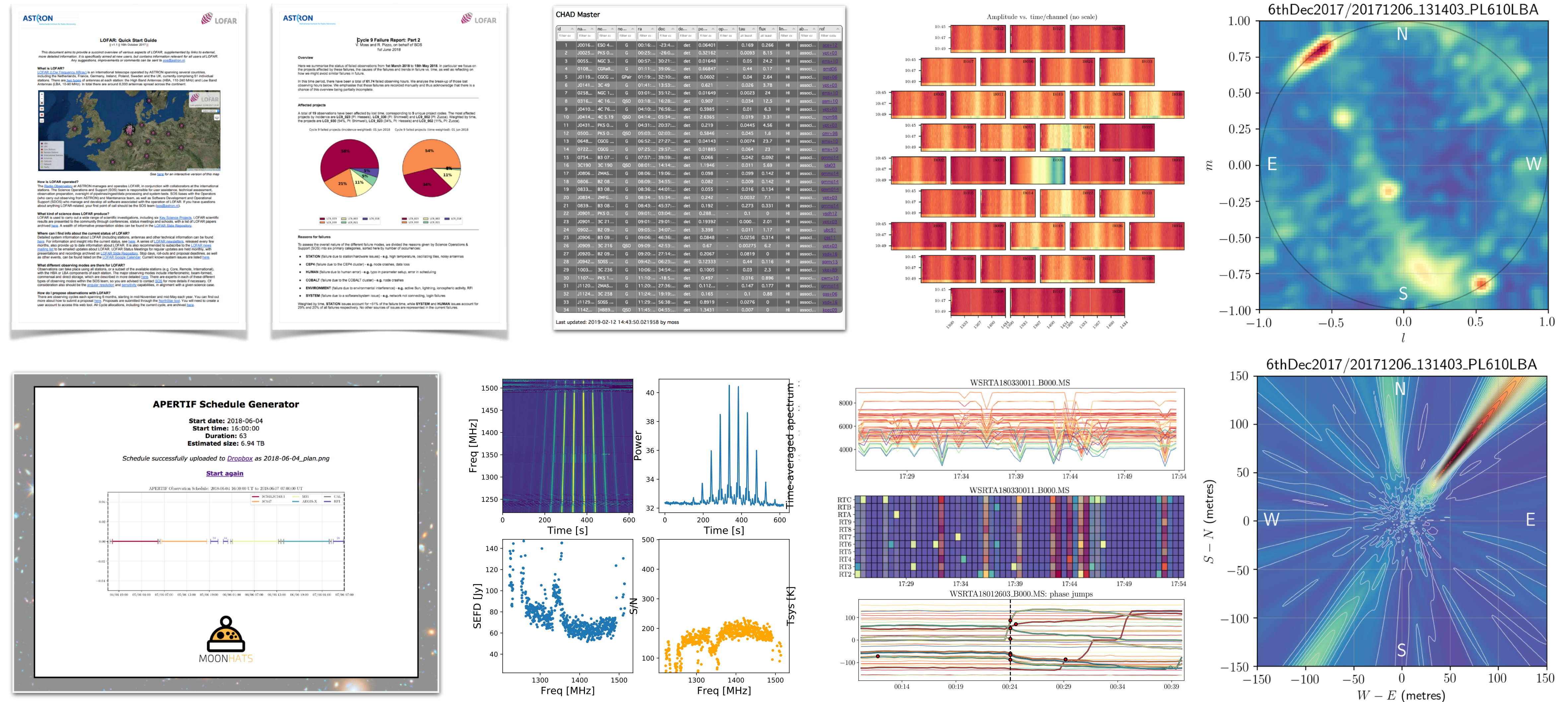


Role of the technical astronomer

- **Observational radio astronomy** has become more specialised and contained in the recent past, compared with the broader experience of past astronomers
- Now, radio astronomers tend to be more focused on specific parts of the overall process, particularly if they are **technically inclined**
- It is particularly important that we **recognise** and **encourage** contributions of technical astronomers in the middle 3 boxes here:



Types of technical contributions



The future of operational roles

- There are many examples of roles at current radio institutes which demonstrate the **growing need** for technical astronomers
- **Science operations:** NRAO support scientists, LOFAR SOS, JIVE telescope scientists
- **Observatory experience:** co-supervision programs, LOFAR traineeships, radio schools
- As a field, we must recognise the need for this role and provide **viable career options** for technical astronomers going forward



2. Automation

Automate all the things

Apercal Triggers APP 18:15

AutoCalBot

AutoCal Status Report: happili-01

Apercal pipeline triggered for 19042805

2019-05-01 16:15:56.830060 UTC

OSA: N/A

FYI: @moss @betsey

Calibrator #1: 3C147 (190428024-190

Calibrator #2: 3C286 (190429001-190

Beams: 0,1,2,3,4,5,6,7,8,9

Show less

AutoCalBot

AutoCal Status Report: happili-04

Apercal pipeline triggered for 19042805

2019-05-01 16:15:57.344727 UTC

OSA: N/A

FYI: @moss @betsey

Calibrator #1: 3C147 (190428054-190

Calibrator #2: 3C286 (190429031-190

Beams: 30,31,32,33,34,35,36,37,38,3

Show less

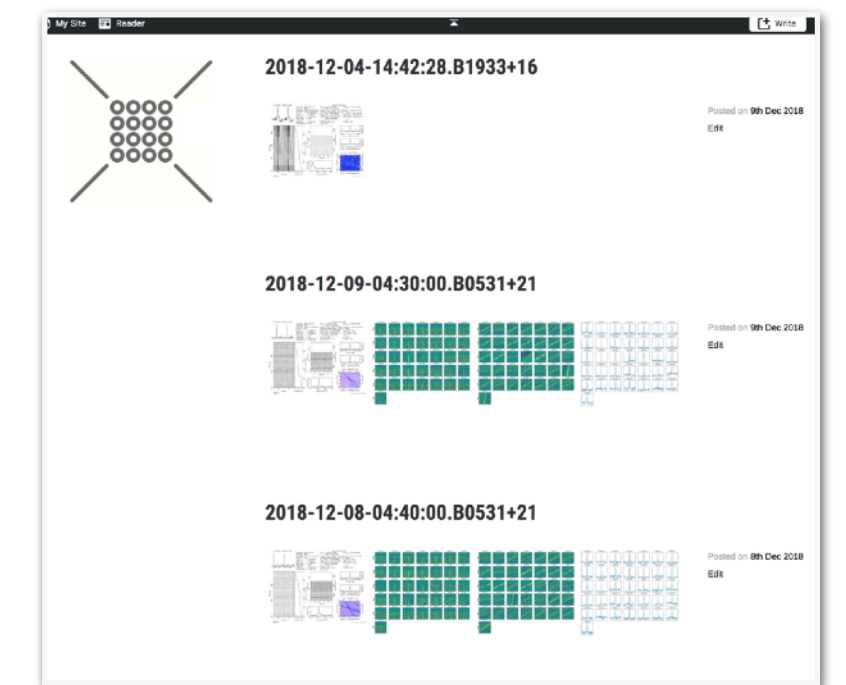
Automating Apertif

- **ATDB:** The Apertif Task DataBase coordinates specification, observations, ingest (ALTA, **P5.1**) and deletion with minimal human effort (**P9.9**)
- **ATDBspec:** Python lib to convert astronomy specification to system specification (+ARTS)
- **Apercal:** a scripted pipeline to process/assess imaging data using preset parameters (**P10.47**), producing science-quality data (**P5.7**)
- **AutoApercal:** auto-detection of new targets
- **ARTS:** real time transient pipeline (**P10.52**)

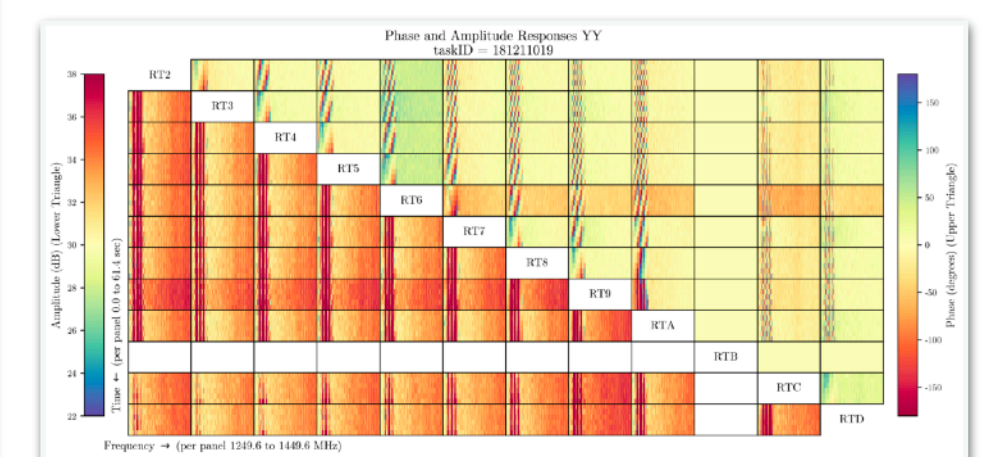
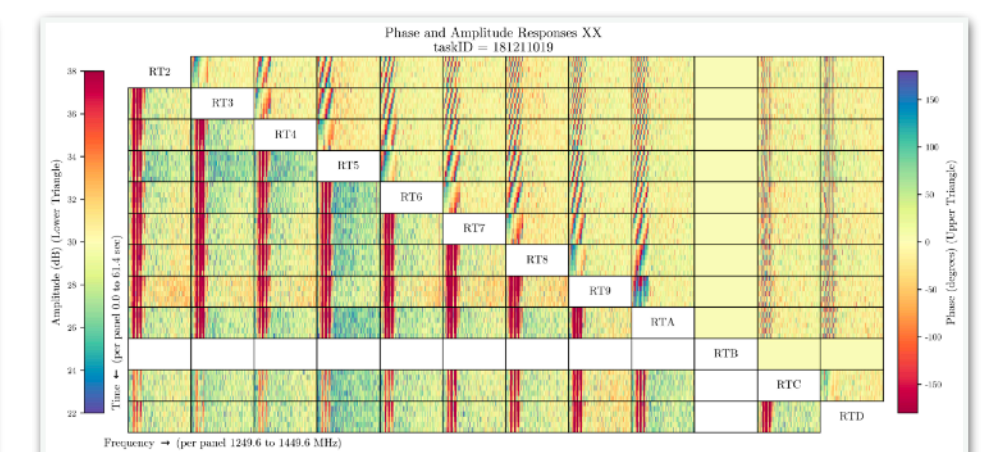
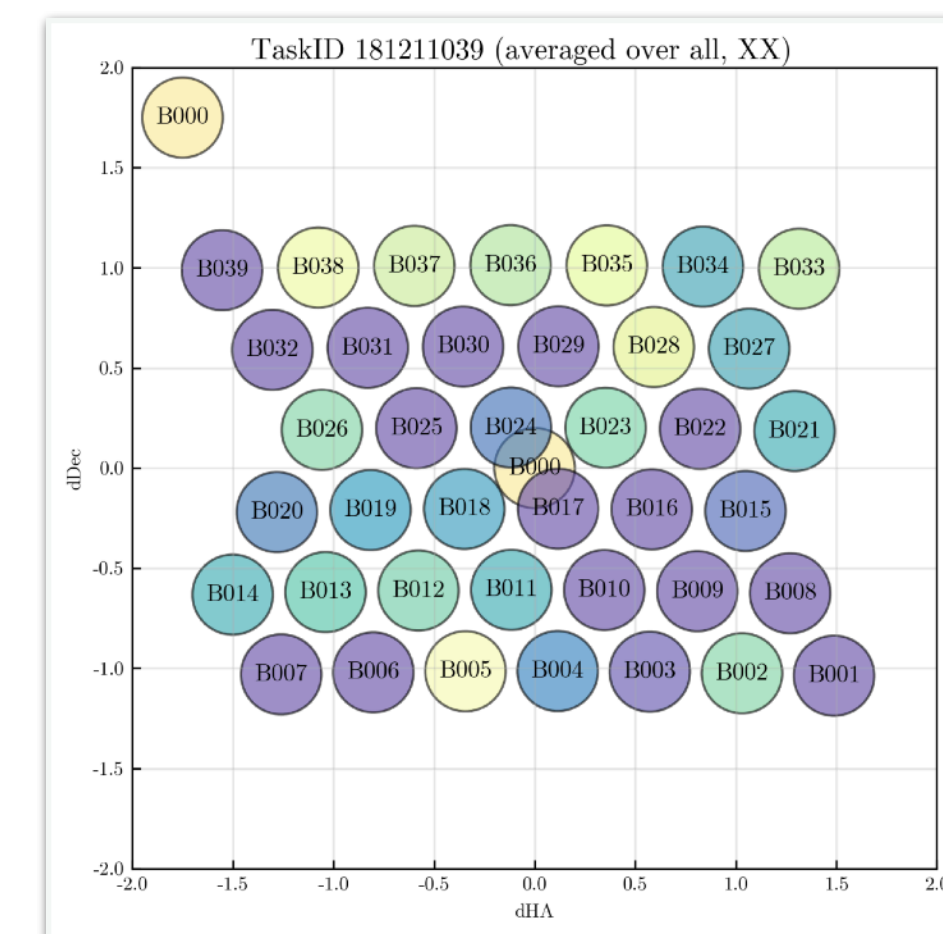
Apertif Task Database

Observations

TaskID	Obs Mode	Name	Status	Start Time	Stop Time
181219004	arts_sc4_survey	B1933+16	starting	2018-12-19 17:00:00	2018-12-19 17:00:00
181219003	arts_sc4_survey	B1933+16	completing	2018-12-19 15:30:00	2018-12-19 15:30:00
181219002	arts_sc4_survey	B1933+16	completing	2018-12-19 15:25:00	2018-12-19 15:25:00
181219001	arts_sc4_survey	B1933+16	completing	2018-12-19 14:07:00	2018-12-19 14:07:00
181218021	arts_sc4_survey	B0531+21	completing	2018-12-18 21:50:00	2018-12-18 21:50:00
181218020	arts_sc4_survey	B0531+21	completing	2018-12-18 21:40:00	2018-12-18 21:40:00

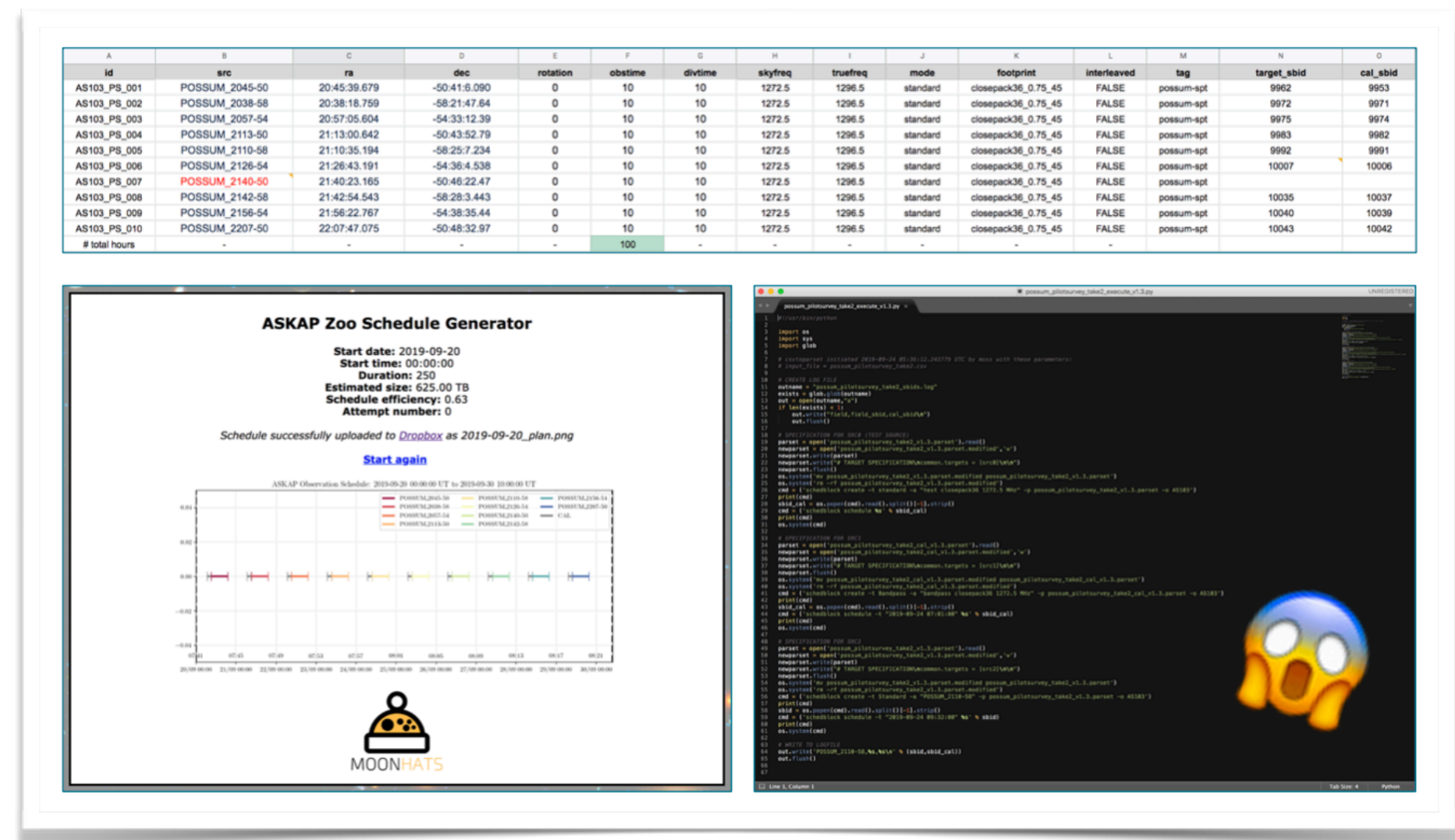


	1	2	3	4
start_pipeline	0:00	0:00	0:00	0:00
prepare	0:26	0:30	0:30	0:26
preflag	7:13	7:16	7:16	10:30
ccal	5:13	5:23	5:10	5:16
QA	0:33	0:16	0:16	0:16
convert	4:30	4:26	4:16	4:30
scal	9:00	6:16	7:43	8:03
continuum	0:00	2:26	1:40	2:43
	26:56	26:36	26:53	31:46



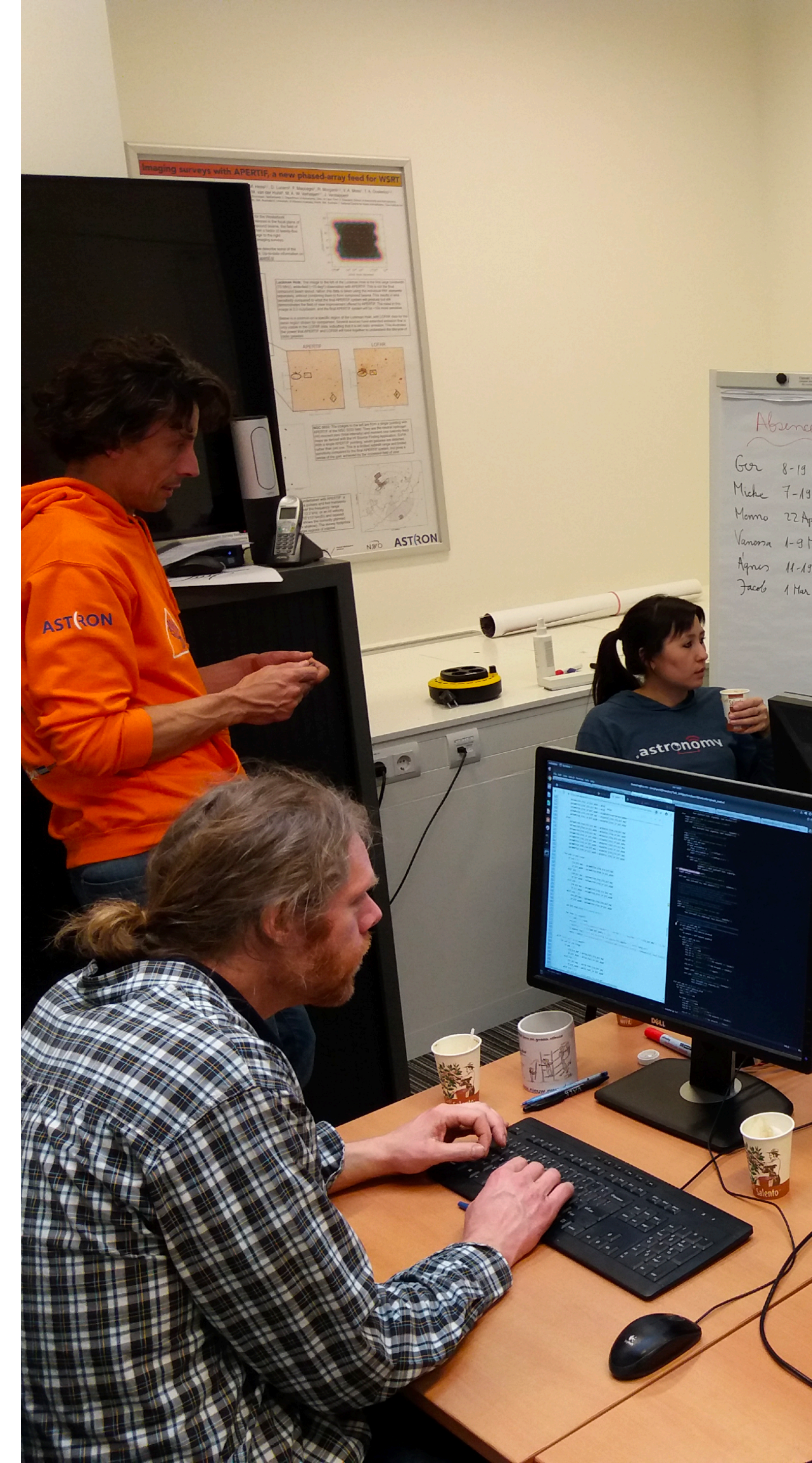
Automating ASKAP

- A lot of work has been done in automating various aspects of ASKAP **prior** to my role (thanks ASKAP team!)
- We make use of an online web tool - **ASKAP OMP** - to provide an **overview** of the system and interact with the **observation database**
- The **ASKAP-X** project has just begun, which is a SAFe Agile software development project to improve **operational stability** and **automation**
- As a scheduler/operator, I can also interact with the system via "**hacky**" Python scripts



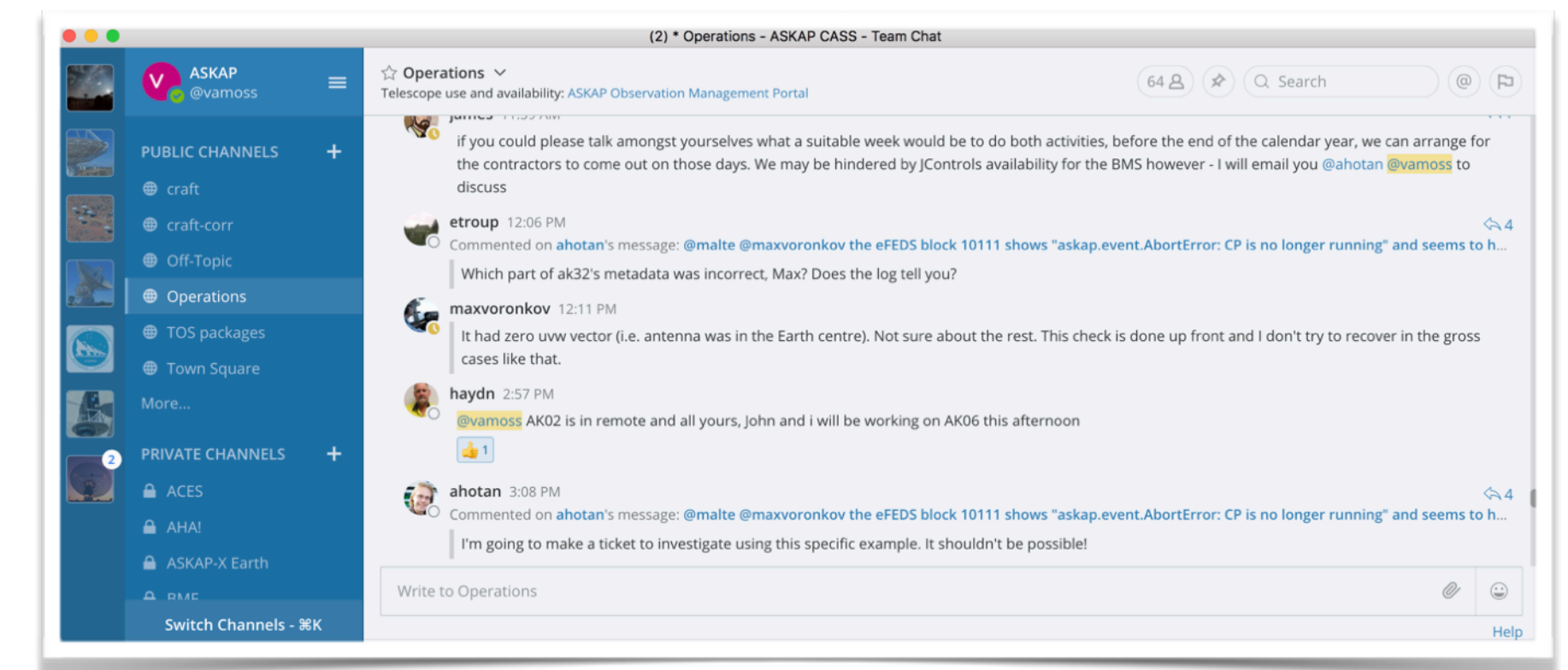
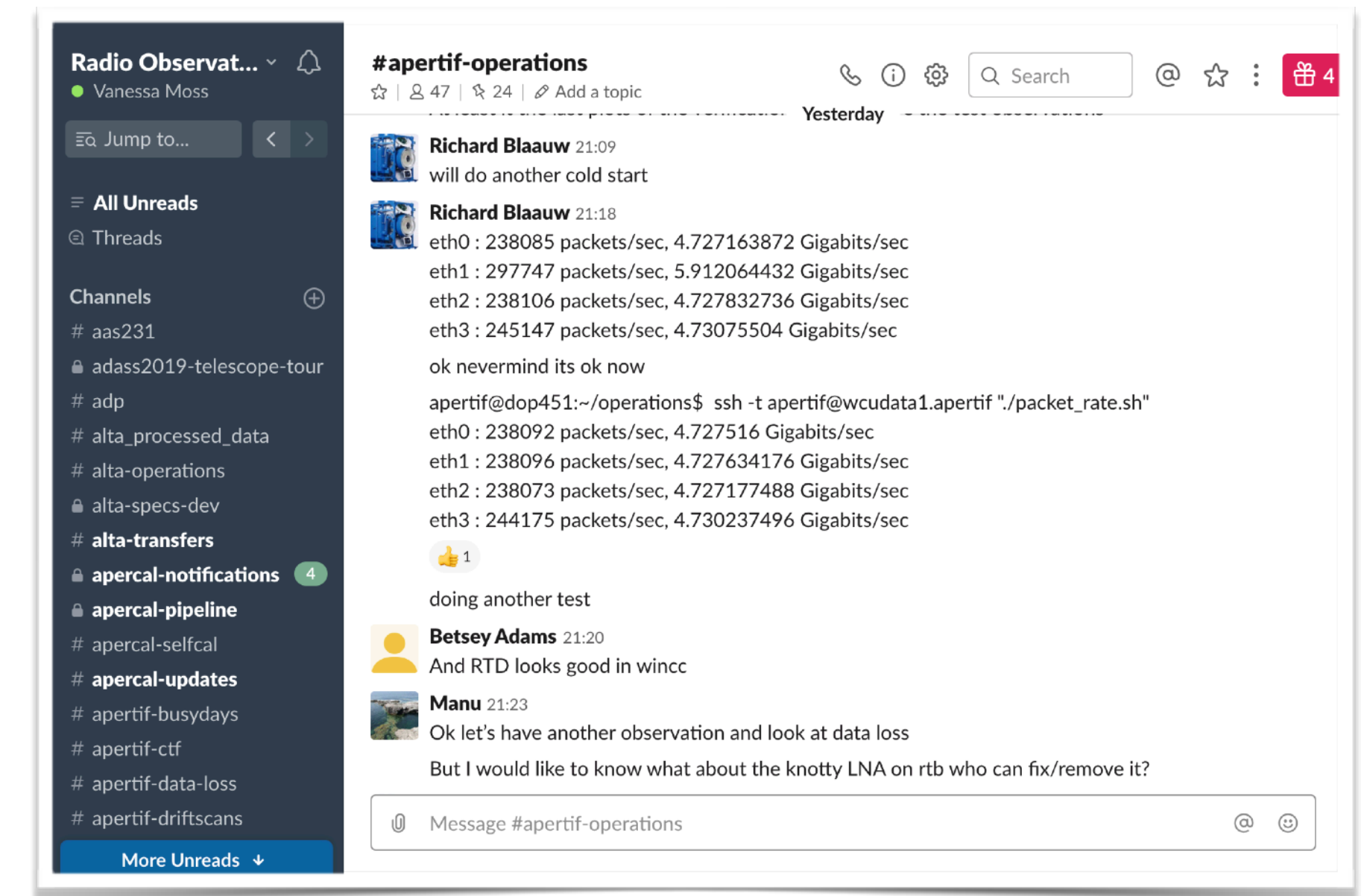
3. Communication

Communication is the source of all the world's problems



The right tool for the right job

- There are **endless ways** to communicate: email, meeting, telecon, phone, **ADASS**, Slack, wikis, reports, WhatsApp, hallway chats, **BoFs!**
note: "dear diary" meetings are rarely useful
- **Daily communication:** Slack or Mattermost are excellent tools to keep the conversation going, especially in **open channels**
- **Weekly communication:** a contained (to the point) meeting, or an online written briefing
- **Long-term communication:** reports or documents, large-scale meetings, telecon



Scheduling meetings

- Meetings can be **powerful**, when used wisely
- **Every Monday 15:30 AEST:** ASKAP scheduling meeting, focused on operations, processing and on-site maintenance (30 MINUTES ONLY!)
- This meeting brings together a very effective group of people to **make a plan** for the week ahead and get everyone on the **same page**
- It is also a very **small tax** on people's time, and we do our best to keep within the time limit
- We notice the **absence** of this meeting!

Pages / ... / 2019 SCD Meetings 3 Jira links

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2019-10-07 Meeting notes

Created by Hotan, Aidan (CASS, Kensington WA), last modified by Moss, Vanessa (CASS, Marsfield) just a moment ago

Date

07 Oct 2019

Attendees

- Apologies from @Hotan, Aidan (CASS, Kensington WA) and @Moss, Vanessa (CASS, Marsfield) who are at ADASS
- Apologies from @Whiting, Matthew (CASS, Marsfield)
- Meeting could be run by @Raja, Wasim (CASS, Marsfield), @Bastholm, Eric (CASS, Kensington WA) or @Hannah, James (CASS, Gerltn Onslow) but we may cancel again as many people are away. Please read these notes!

Goals

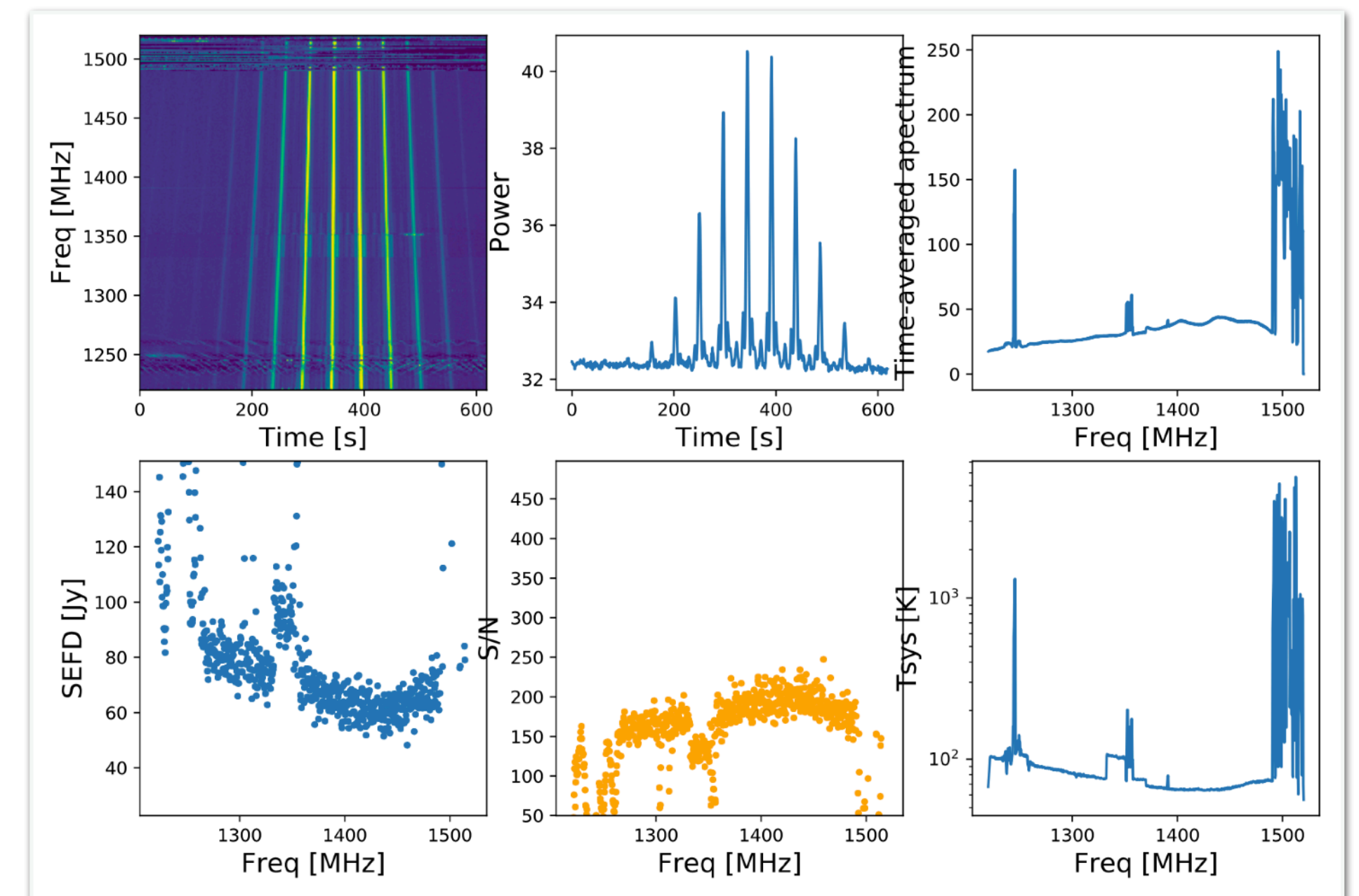
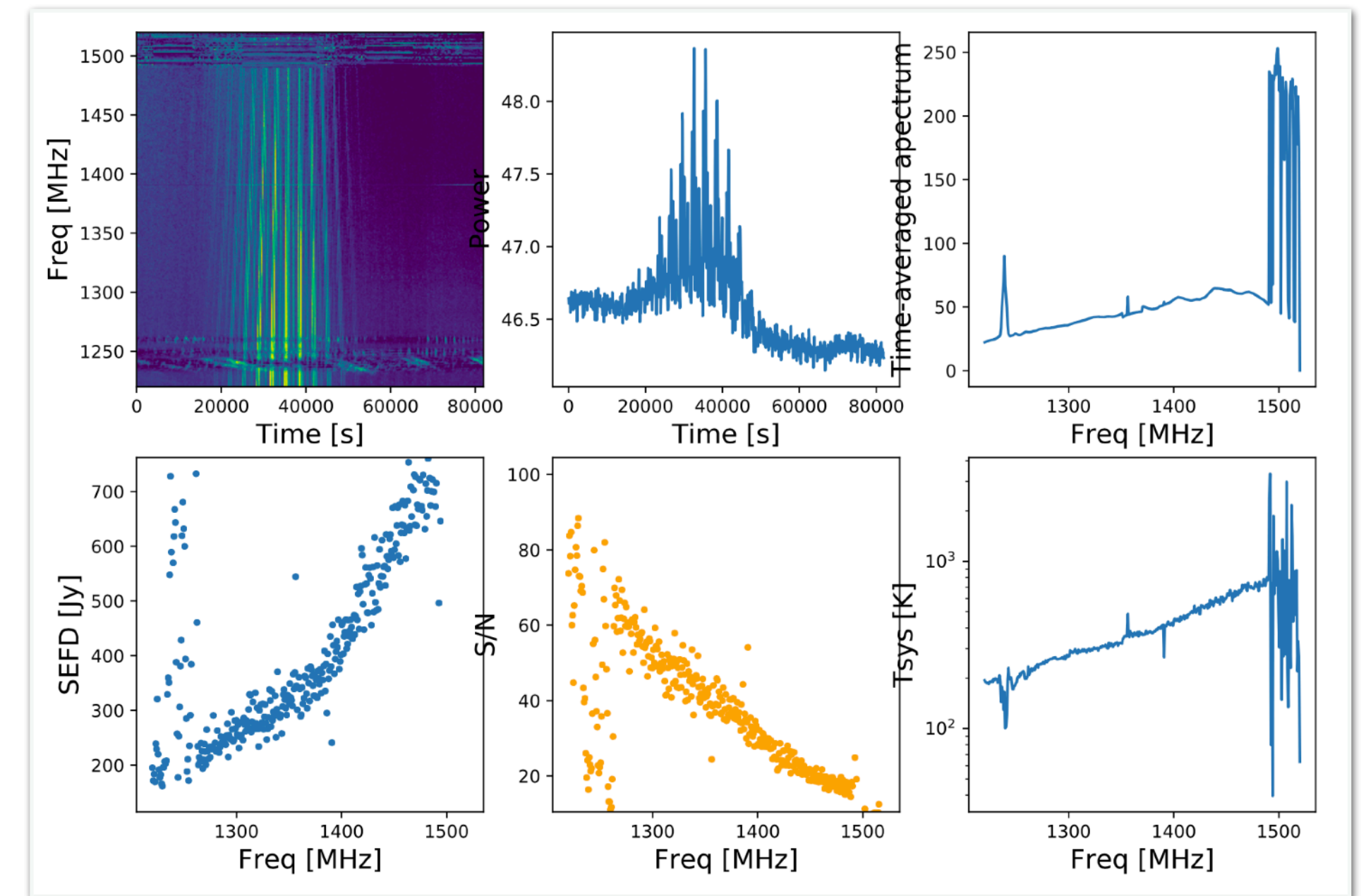
- Set scheduling priorities and resource allocation for the week

Discussion items

Time	Item	Who	Notes
10 min	Telescope schedule	Aidan/Vanessa	<ul style="list-style-type: none">• eFEDS test fields over the weekend• First priority over the next two weeks is deployment and full-array testing of the new array sync functionality developed by AXE team<ul style="list-style-type: none">• Awaiting packages suitable for installation and acceptance testing• Second priority for the next two weeks should be on the integration of the new disk array at Pawsey<ul style="list-style-type: none">• This is now partitioned with two mount points: /askapbuffer/ingest and /askapbuffer/processing• Transfer script ticket: ASKAPSDP-3712 - New disk buffer: define cp process IN PROGRESS• Deployment ticket: ASKAPSDP-3565 - Deployment of new disk buffer OPEN• Proposed steps:<ol style="list-style-type: none">1) Switch the ingest pipeline to write to /askapbuffer/ingest2) Develop scripts to copy observations from ingest to processing and delete from ingest3) Switch the data processing pipeline to read from processing4) Copy everything from /astro over to /askapbuffer/processing (we could trim this a bit, but we need to keep all the unprocessed pilot data at the very least)• Pilot Survey update:<ul style="list-style-type: none">• EMU: 10/10 done• CRAFT: 7/11 done (but the rest are postponed)• WALLABY: 1/3 done (but field probably needs reobserving due to dropouts)• VAST: 1/6 done (pending processing)• POSSUM: 10/10 done (pending processing)• GASKAP: N/A• FLASH: N/A• DINGO: N/A• Observing plan: we need to get back to spectral line observations given the current correlator stability<ul style="list-style-type: none">• This depends on disk space (so we cannot proceed until the new disks are integrated)• 1) GASKAP test field• 2) FLASH subset of pilot survey• 3) WALLABY field 2/3 Norma is suitable• 4) DINGO ???• In the meantime, due to not all dishes being available + disks pending, we will continue with eFEDS continuum tests
10 min	Pawsey schedule	Wasim	<ul style="list-style-type: none">• WALLABY Eridanus processing now finished? If not, how much more to do?• Next priority is SMC and LMC processing according to GASKAP specifications as supplied here: GASKAP• Need to slot the repeated EMU fields in around these other activities• Would be good to investigate<ul style="list-style-type: none">• ASKAPSUP-1026 - Inconsistent delay across the frequency band on two antennas IN PROGRESS since this seems to be a new failure mode?
10 min	MRO schedule	James/Brett	<ul style="list-style-type: none">• Tues: AK02 (@haydn), AK06 (@johnmorris)• Wed: AK13 (@haydn), AK15 (@johnmorris)• Thu: AK13 (@haydn), AK31 (@johnmorris)• All antennas will be returned to the array at the end of each day (4pm AWST).

Results from ACTF

- **ACTF:** Apertif Calibration Task Force, initiated by A. Schoenmakers, led by V. Moss/B. Adams
- Consisted of members from across the ASTRON departments to **consult** on calibration issues, including A. Hotan from CSIRO/ASKAP
- 4+ meetings (**27/3, 18/4, 21/5, 18/6**) to address current issues and suggest paths to solutions
- **Outcomes:** warm start + beam weights successfully improved sensitivity, progress in debugging cause of off-beam phase slopes

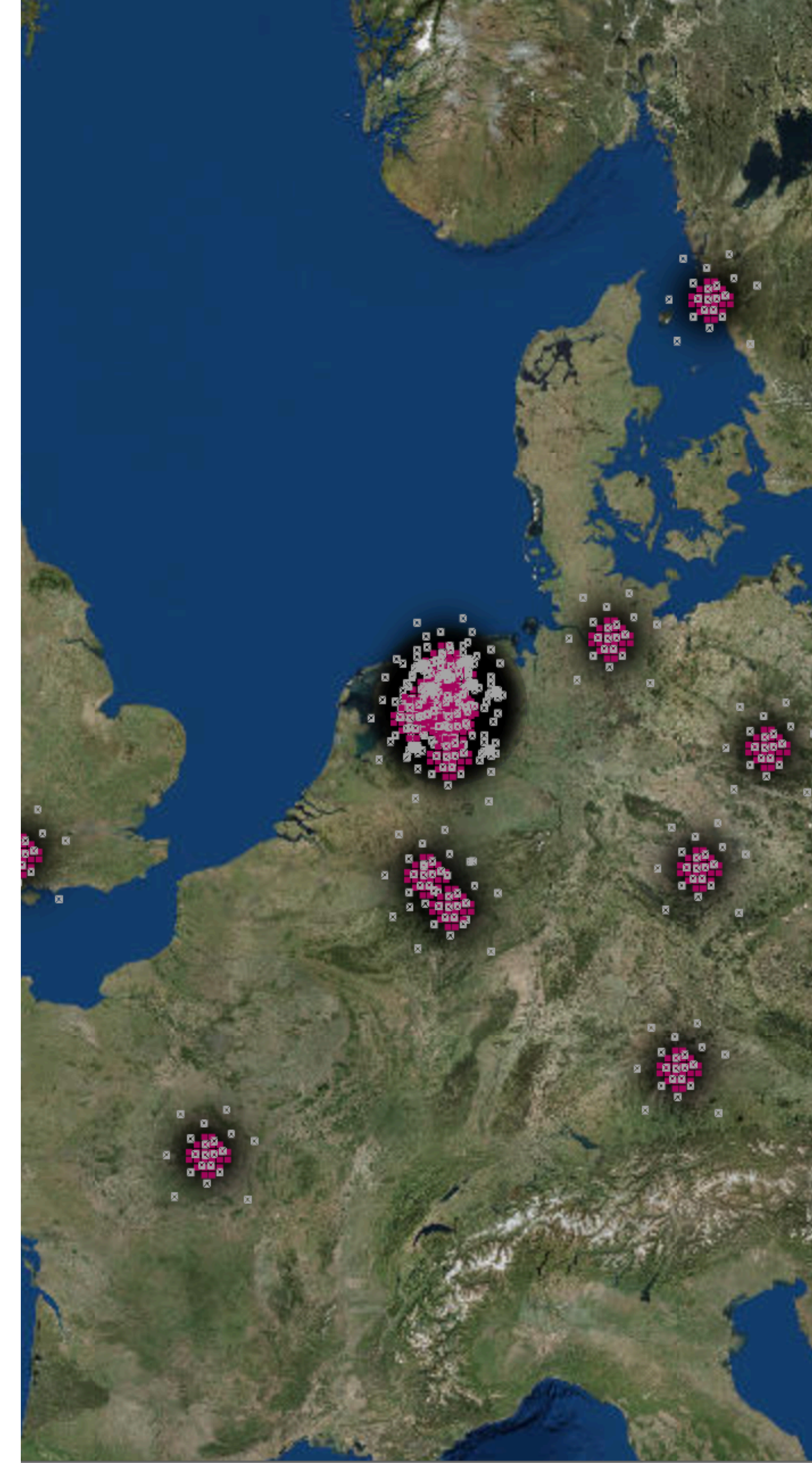


ADASS BoF #4

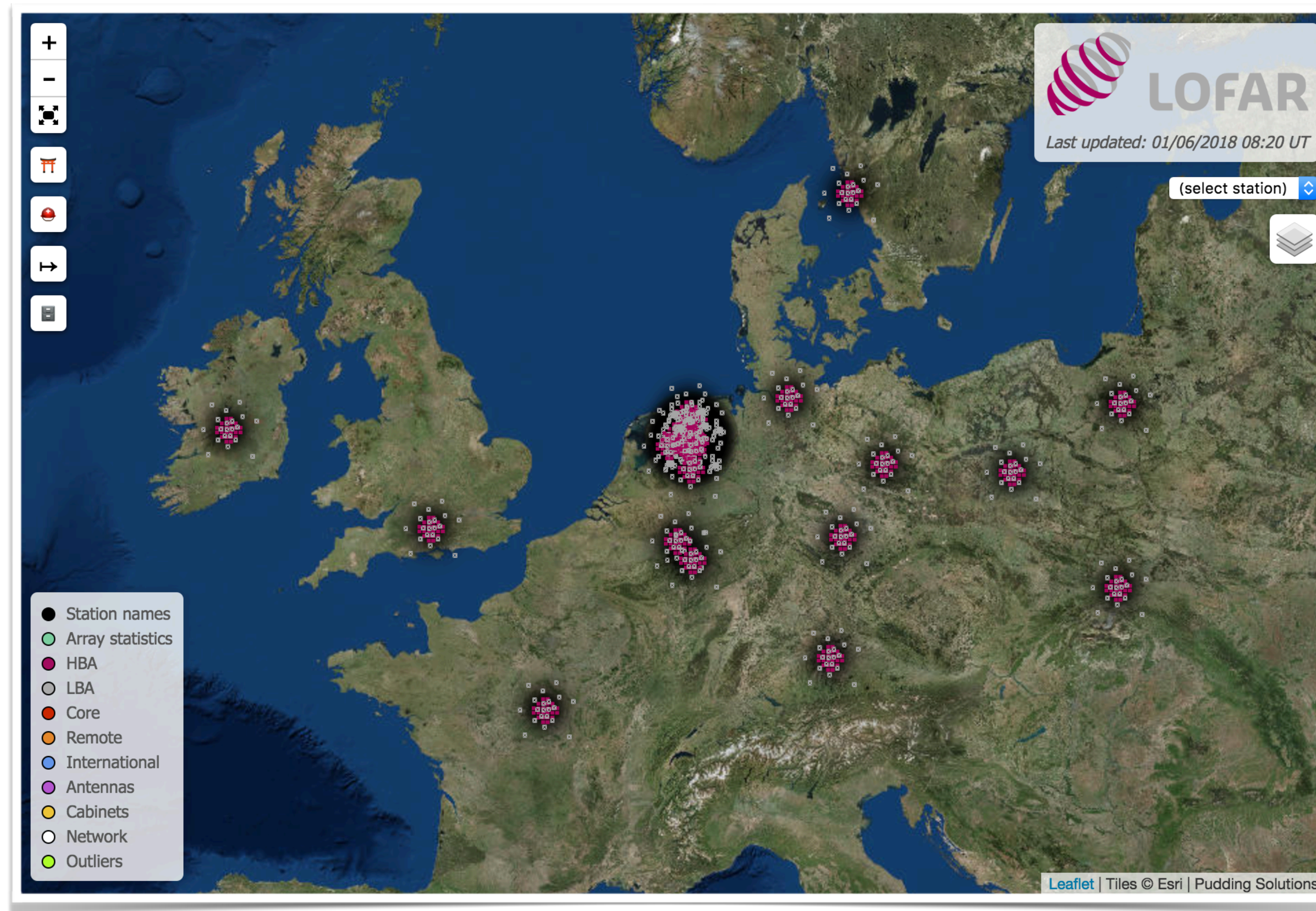


4. Overview

The big picture is just as important as the details



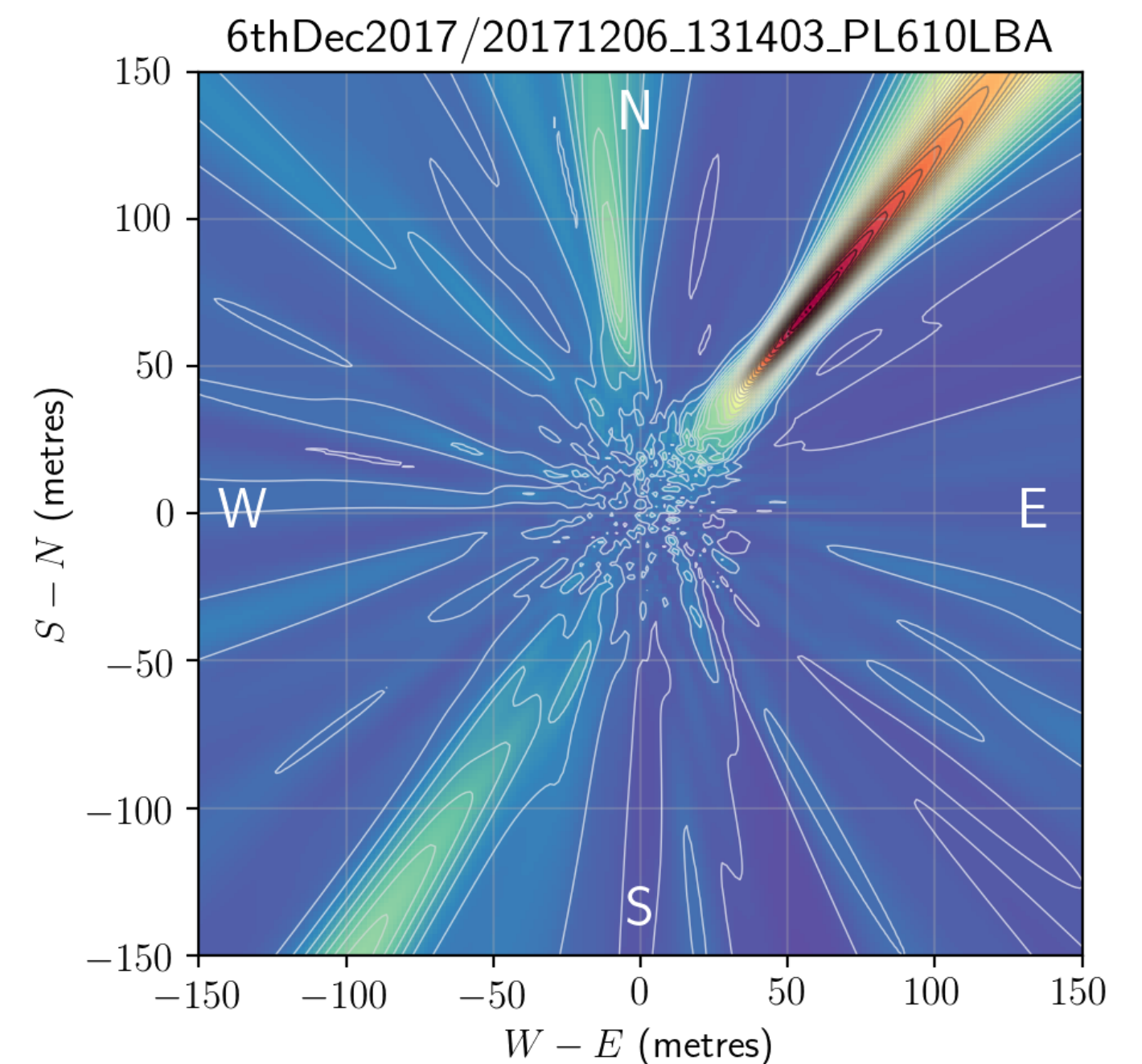
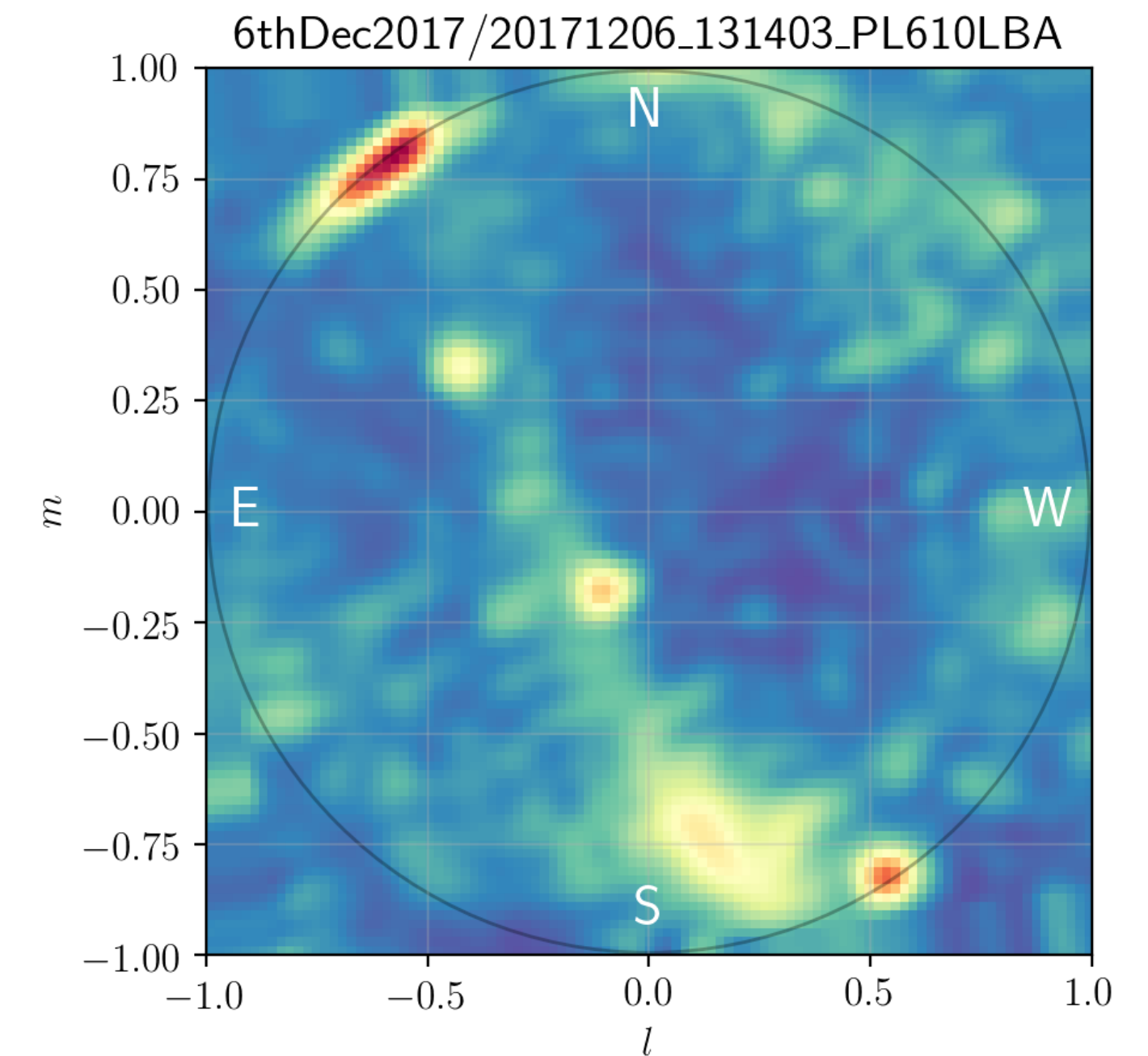
Interactive LOFAR map



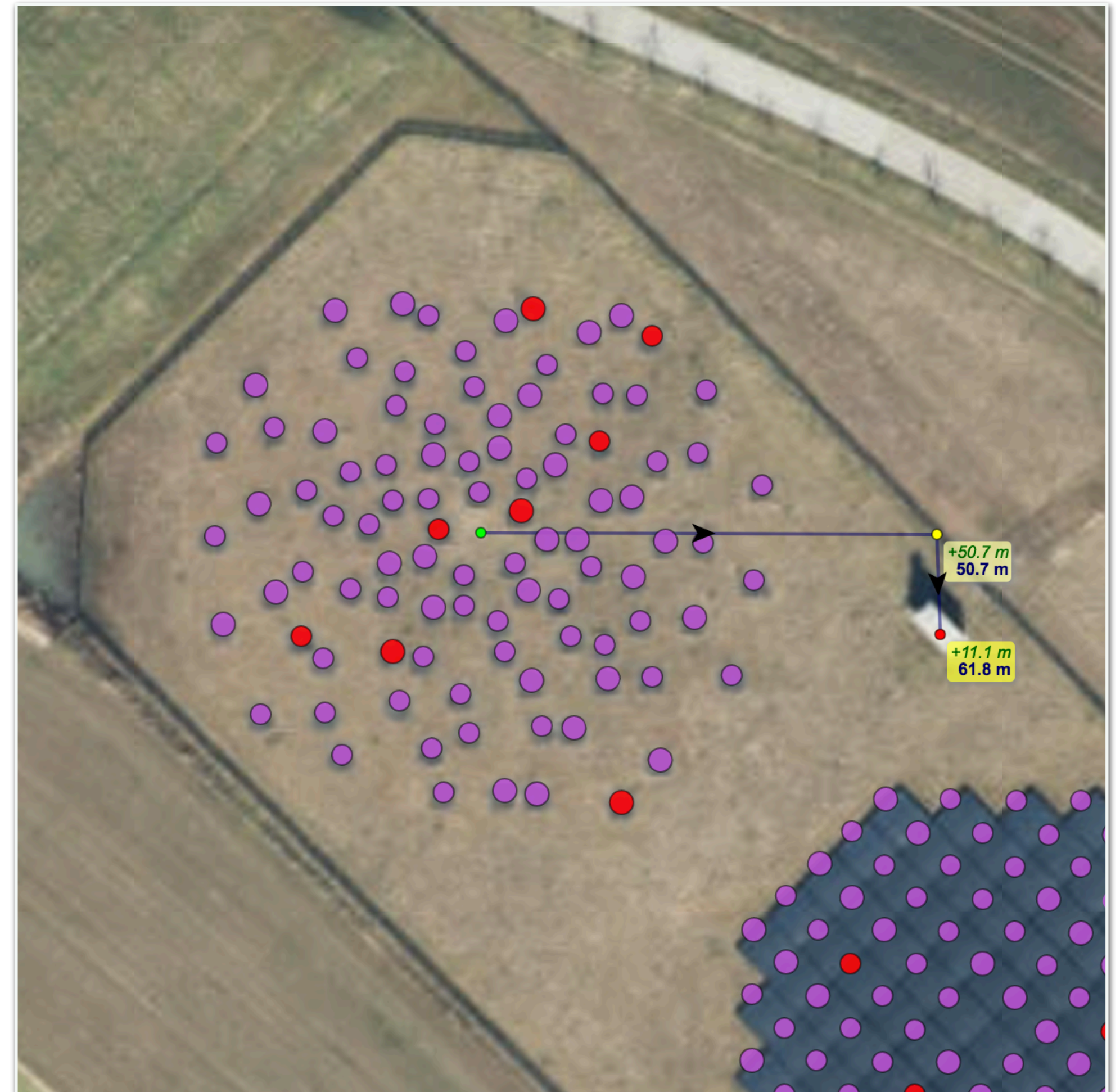
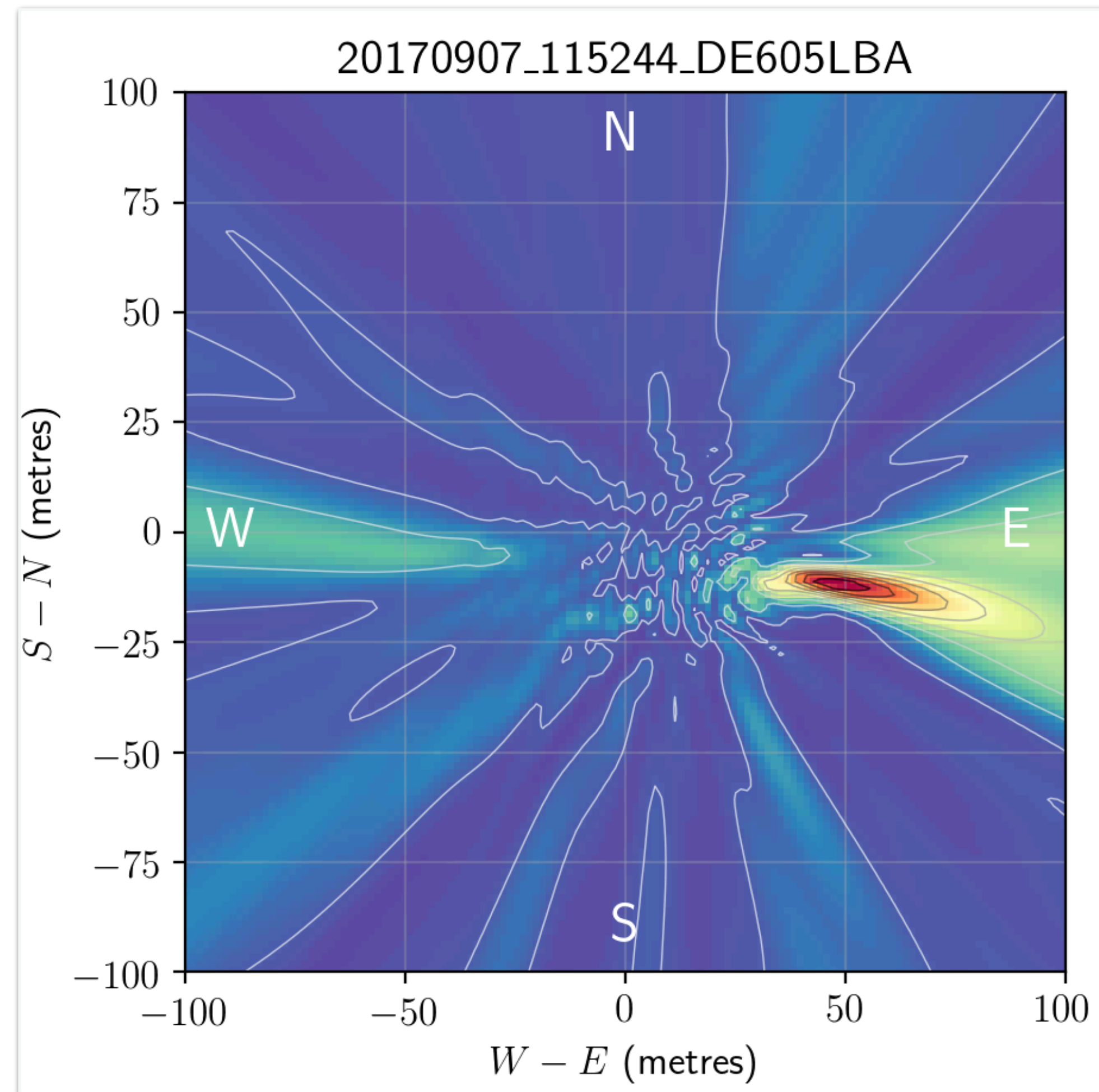
<http://astron.nl/lofartools/lofarmap.html>

RFI near-field imaging

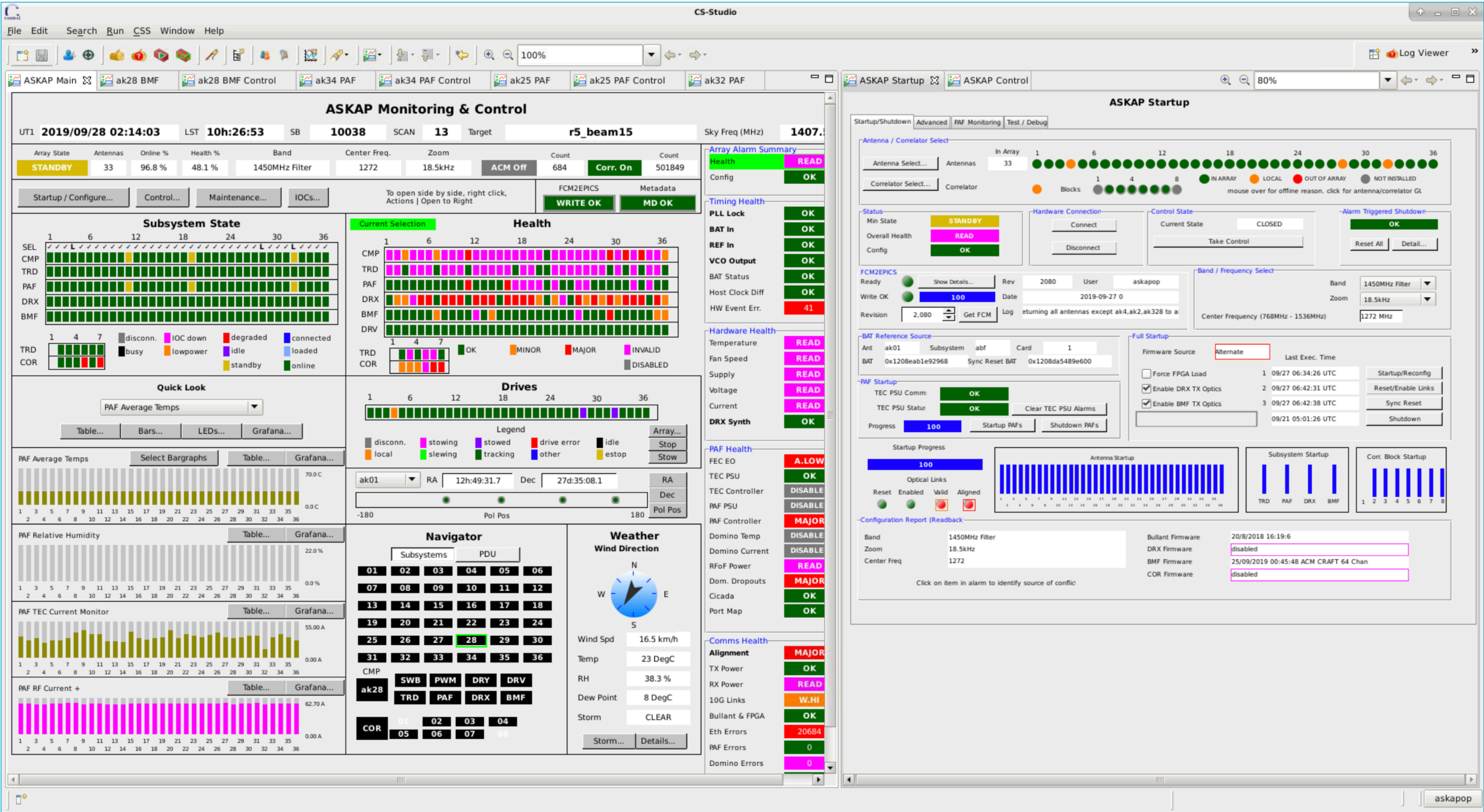
- Began as a tutorial on **quick imaging** of LOFAR data from M. Brentjens
- We record data from an individual station, in the form of the **cross-correlation matrix** and then image this using Python
- From this, it is possible to image **both the sky and the ground plane** - if there is RFI, we can hunt for its location
- Alongside detections of RFI, this method gives us insight into the **environmental overview**



RFI near-field imaging

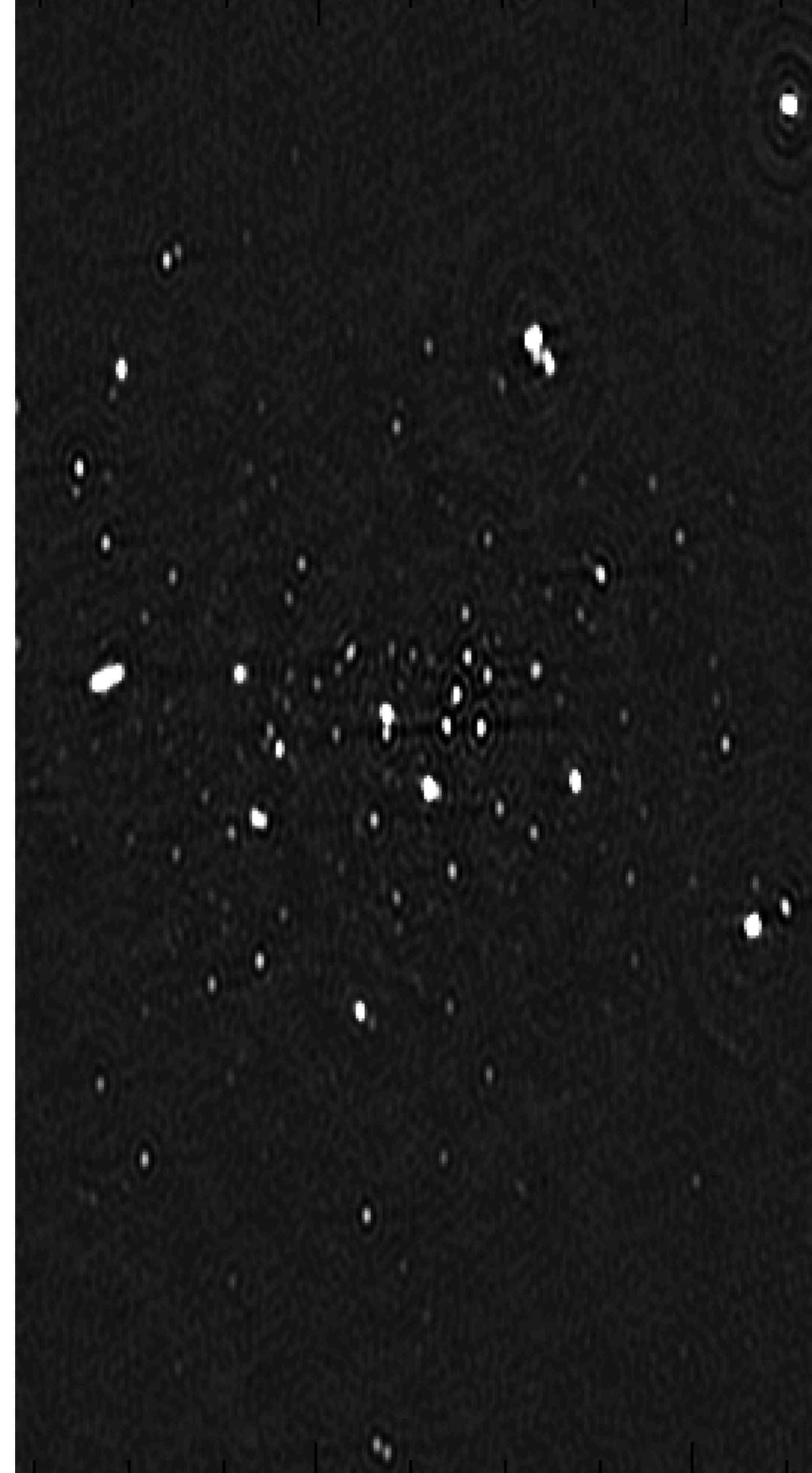


ASKAP CSS



5. Testing

There is no test like an observational test



The SVC


- **SVC:** Science Verification Campaign
- A month-long campaign instigated by the MT to focus specifically on the **data quality** for the imaging and time-domain Apertif surveys
- **All development halted:** observations and calibration only to take place on system
- Imaging and ARTS team then conducted observations in as close to **survey mode** as possible, with data quality to be quantified
- Precursored by 2 operational "**shakedowns**"

APERTIF shakedown month							
Week of: maart 18							
	18/3 MAANDAG	19/3 DINSdag	20/3 WOENSDAG	21/3 DONDERDAG	22/3 VRIJDAG	23/3 ZATERDAG	24/3 ZONDAG
8:00 AM							
8:30 AM							
9:00 AM							
9:30 AM	Coordination						
10:00 AM							
10:30 AM	Phase tuning tests (offset beam delay calibration)						
11:00 AM							
11:30 AM							
12:00 PM							
12:30 PM							
1:00 PM	40 beam tuning 6 bit mode delays + phases square_39p1	ARTS SC4	ARTS SC4	ARTS SC4	ARTS SC4	ARTS SC4	ARTS SC4
1:30 PM							
2:00 PM							
2:30 PM							
3:00 PM							
3:30 PM	40 beam cal scans 6-bit mode 1 min per beam						
4:00 PM							
4:30 PM							
5:00 PM	ARTS SC4						
5:30 PM							
6:00 PM							
6:30 PM							

APERTIF shakedown month							
Week of: maart 25							
	25/3 MAANDAG	26/3 DINSdag	27/3 WOENSDAG	28/3 DONDERDAG	29/3 VRIJDAG	30/3 ZATERDAG	31/3 ZONDAG
8:00 AM							
8:30 AM	ARTS SC4						
9:00 AM							
9:30 AM	Coordination						
10:00 AM							
10:30 AM	40 beam cal scans 6-bit or 8-bit mode 1 min per beam						
11:00 AM							
11:30 AM							
12:00 PM							
12:30 PM							
1:00 PM	40 beam tuning 6-bit or 8-bit mode delays + phases square_39p1	IMAGING	IMAGING	IMAGING	IMAGING	IMAGING	IMAGING
1:30 PM							
2:00 PM							
2:30 PM							
3:00 PM							
3:30 PM	40 beam cal scans 6-bit or 8-bit mode 1 min per beam						
4:00 PM							
4:30 PM							
5:00 PM							
5:30 PM	IMAGING						
6:00 PM							
6:30 PM							

SVC highlights



 **Apercal Triggers** APP 18:15

AutoCalBot

AutoCal Status Report: happili-01

Apercal pipeline triggered for 190428055: LH_WSRT
2019-05-01 16:15:56.830060 UTC

OSA: N/A

FYI: @moss @betsey

Calibrator #1: 3C147 (190428024-190428015)

Calibrator #2: 3C286 (190429001-190429010)

Beams: 0,1,2,3,4,5,6,7,8,9

Show less

AutoCalBot

AutoCal Status Report: happili-04

Apercal pipeline triggered for 190428055: LH_WSRT
2019-05-01 16:15:57.344727 UTC

OSA: N/A

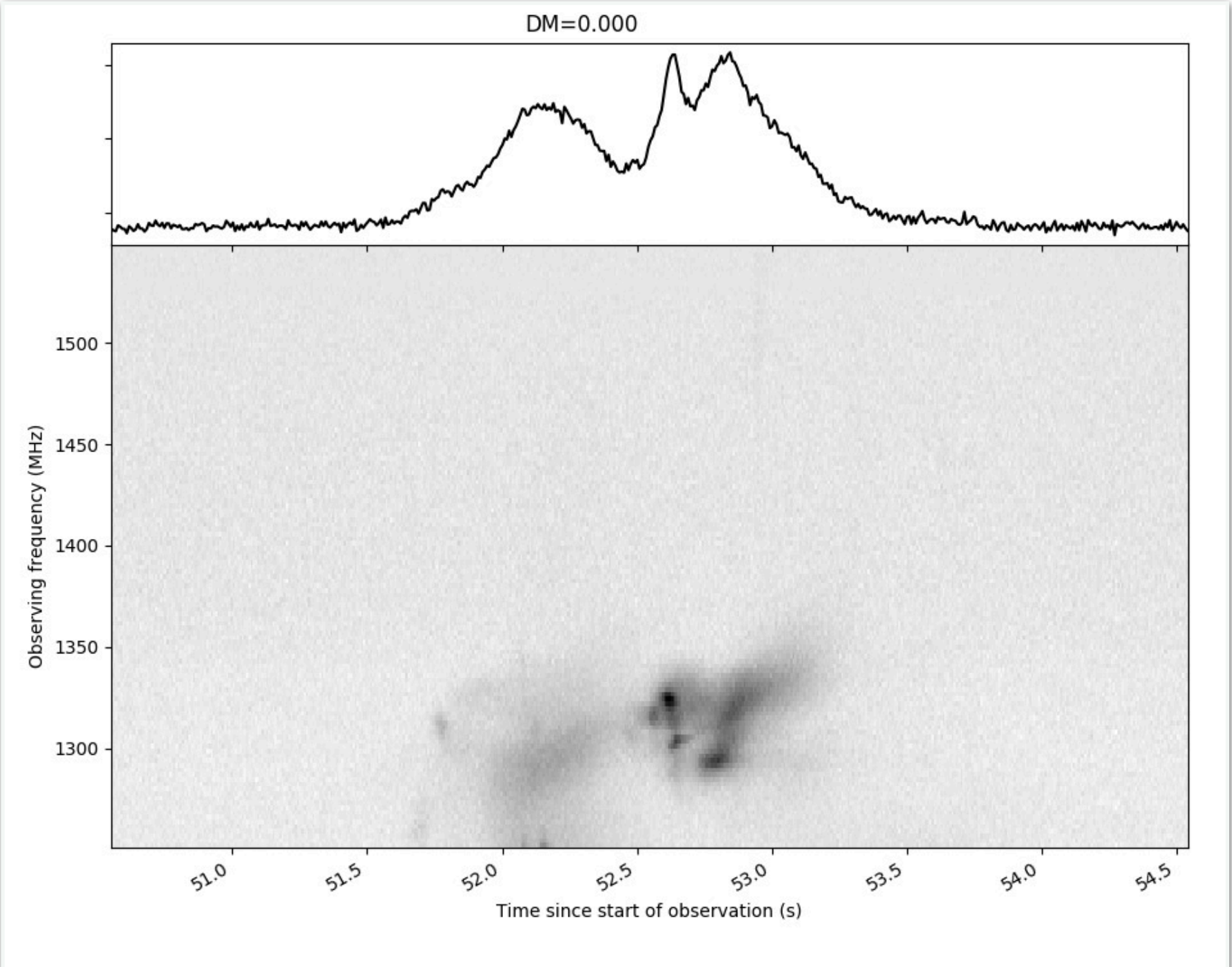
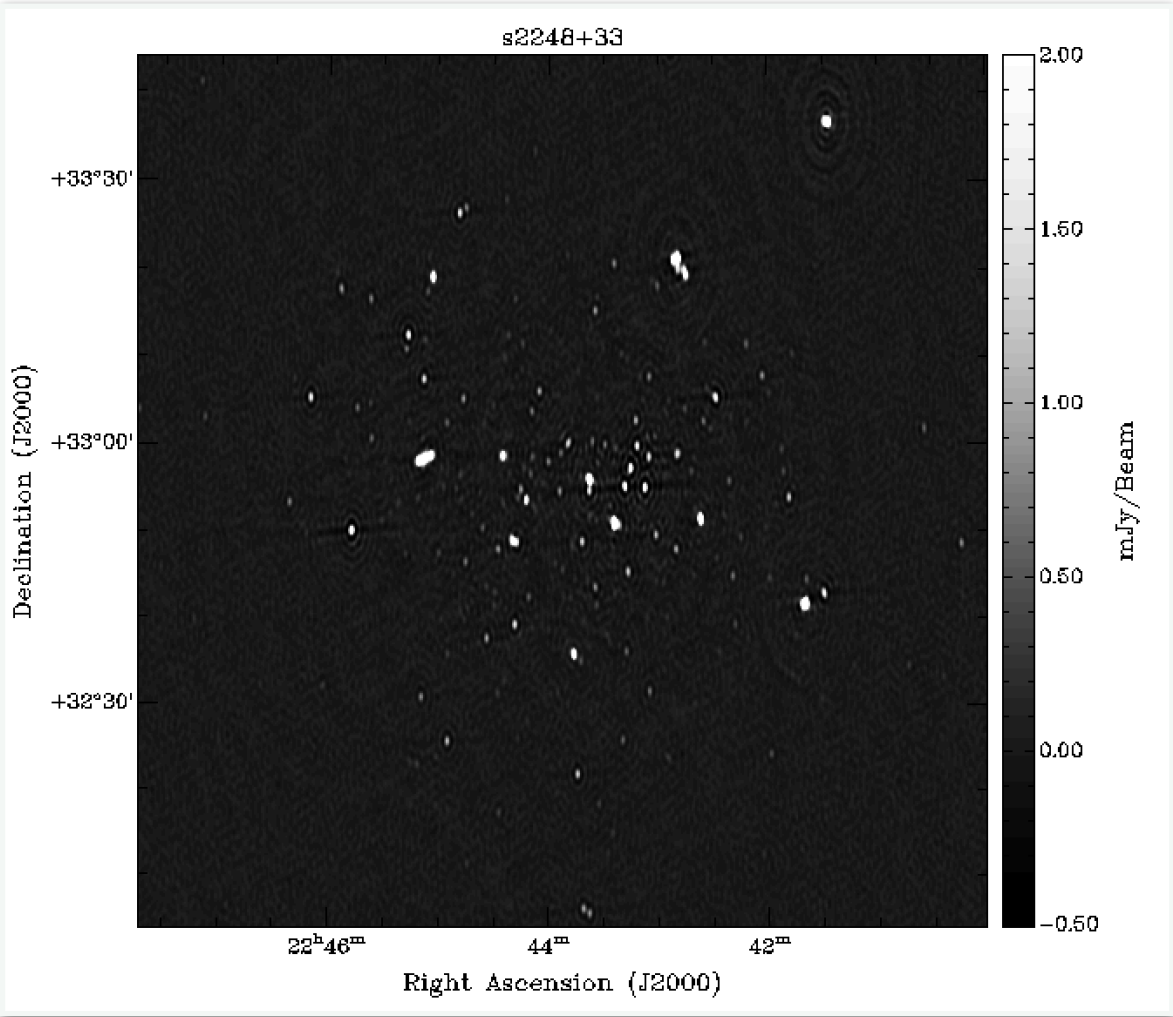
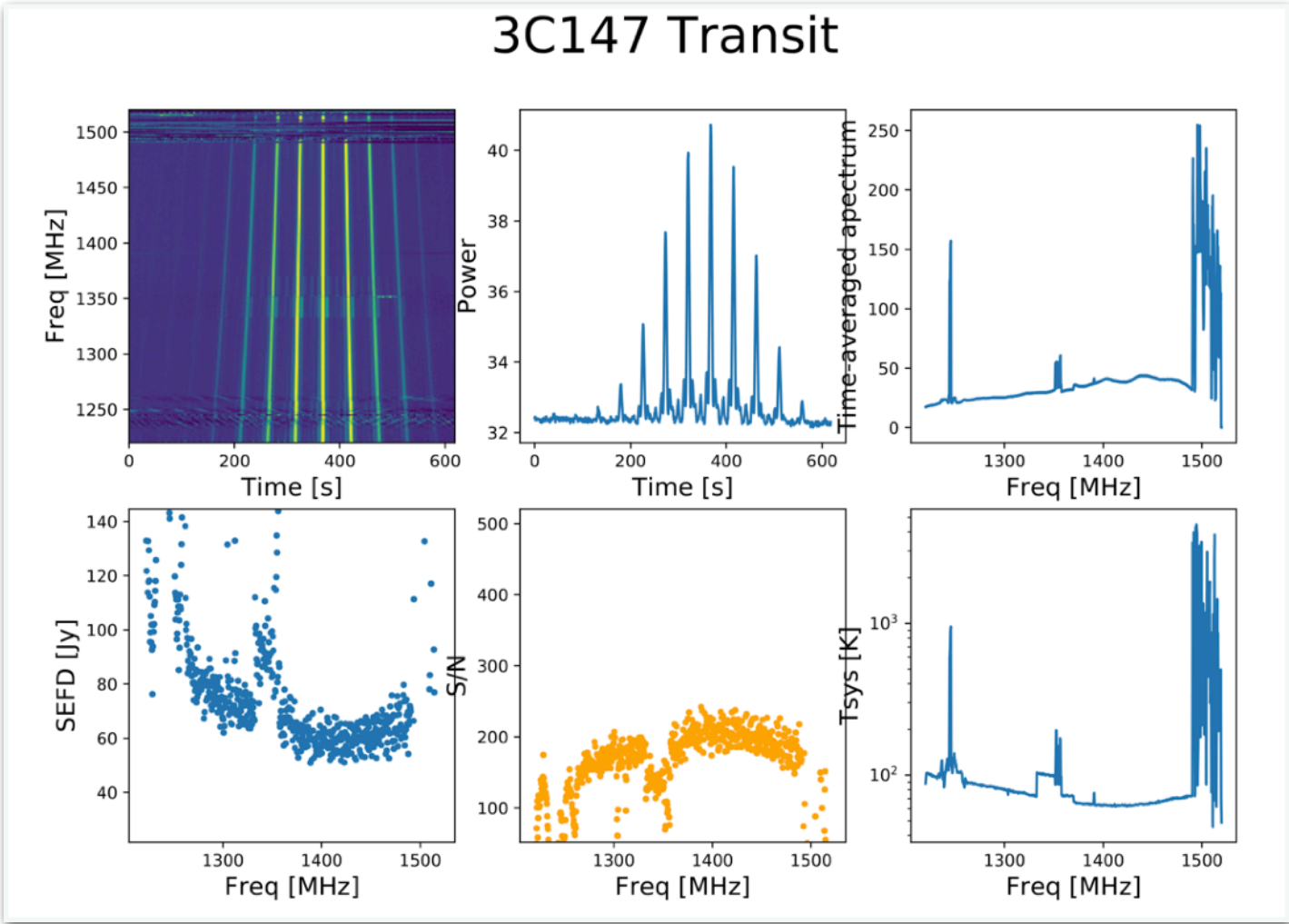
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Calibrator #1: 3C147 (190428054-190428045)

Calibrator #2: 3C286 (190429031-190429040)

Beams: 30,31,32,33,34,35,36,37,38,39

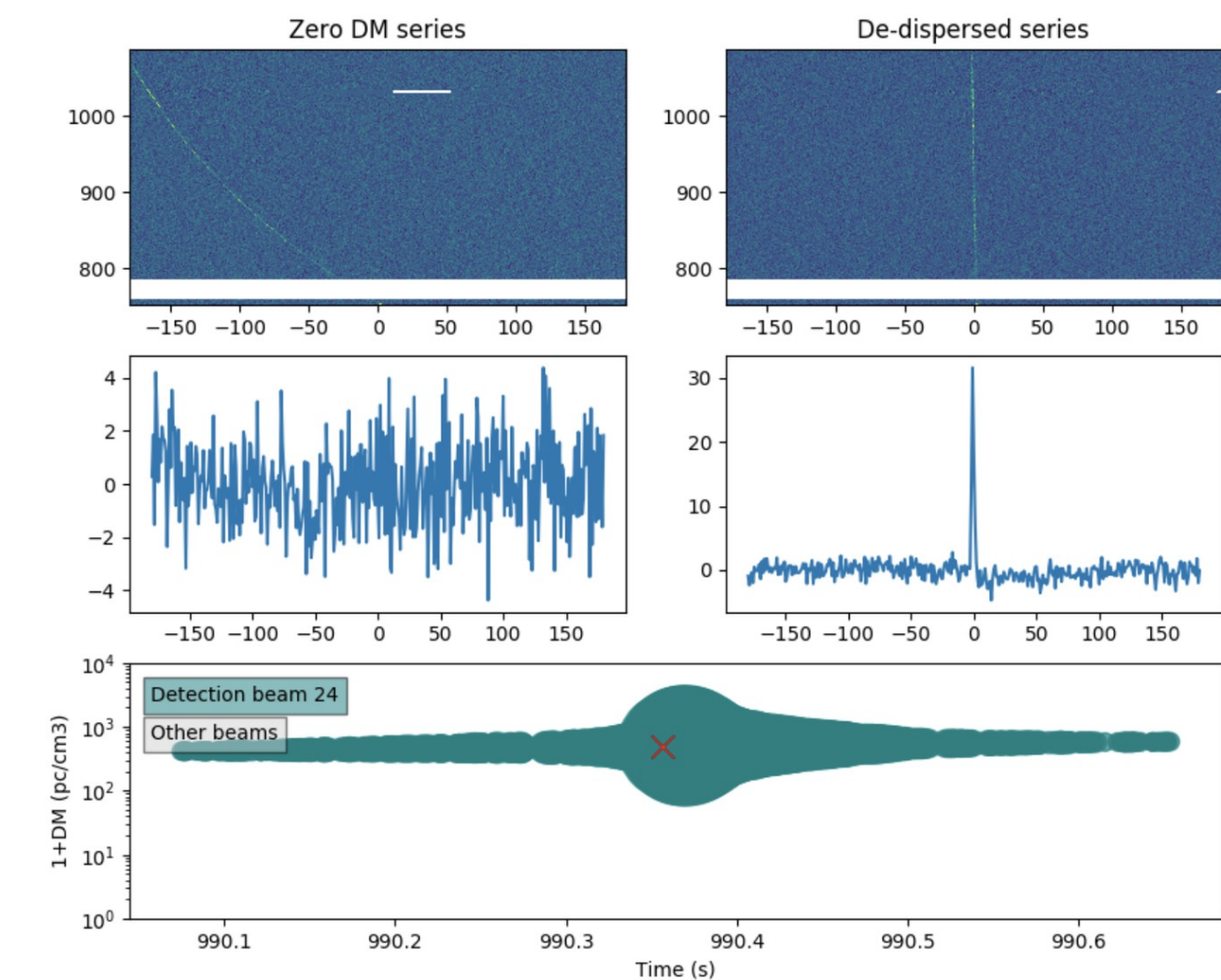
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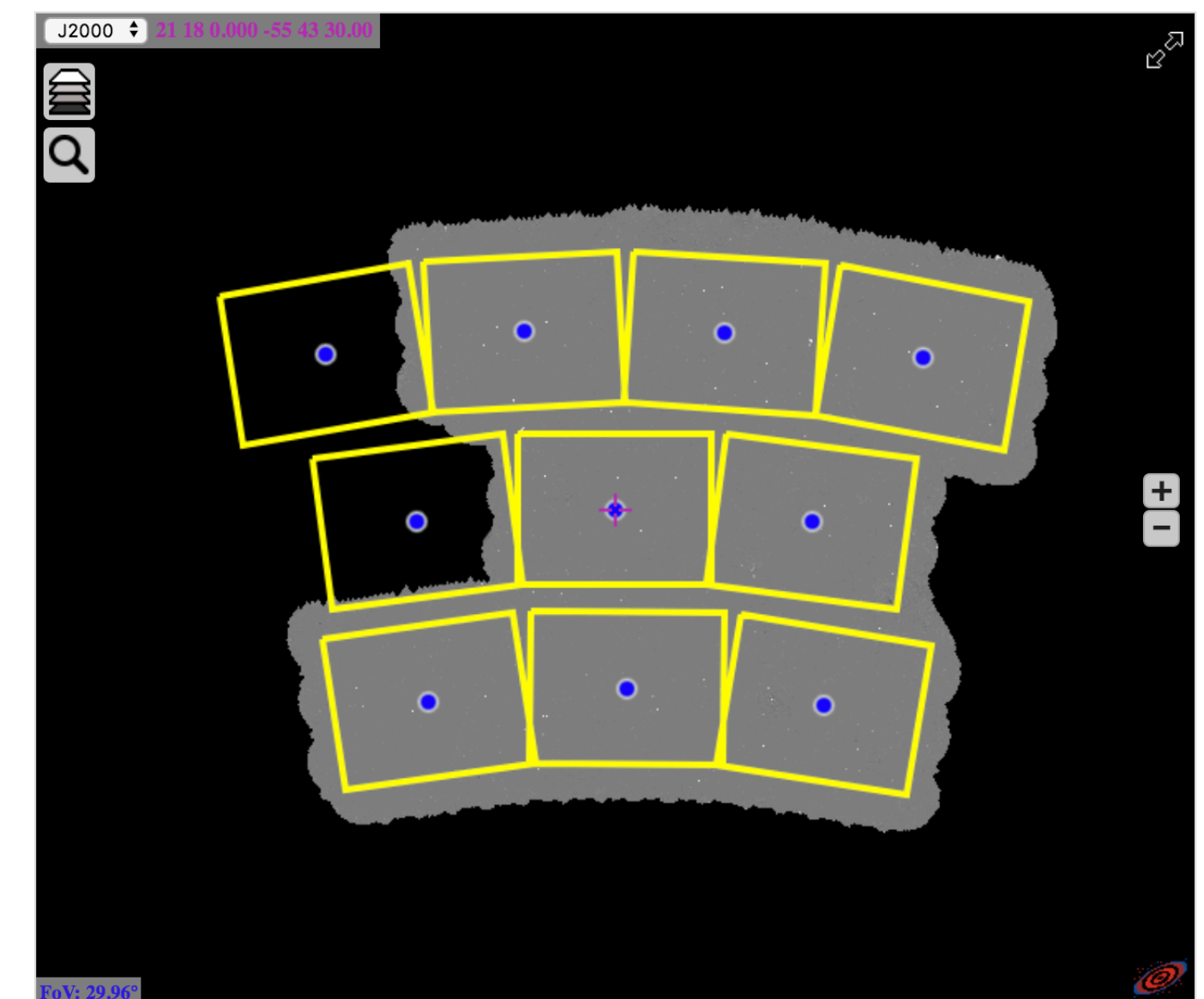
ASKAP Pilot Surveys

- Pilot Surveys are **100 hr** blocks of time allocated to existing ASKAP SSTs
- The Pilot Surveys have a **number of goals**:
 - Give each SST a **representative** set of data to work with
 - Test full survey strategy on a **subset** of sky
 - Push system to **operational limits** to find pressure points
 - Use Pilot Survey period to inform operations of **full surveys**
 - Assess ASKAP pipeline **processing readiness (P10.22)**
- We started the Pilot Surveys on **15th July 2019**
- EMU, WALLABY (1/3), VAST (1/6), CRAFT (7/10), POSSUM now “**complete**”! More to come...

The Great Radio School FRB of 2019



EMU 940 MHz Pilot Survey



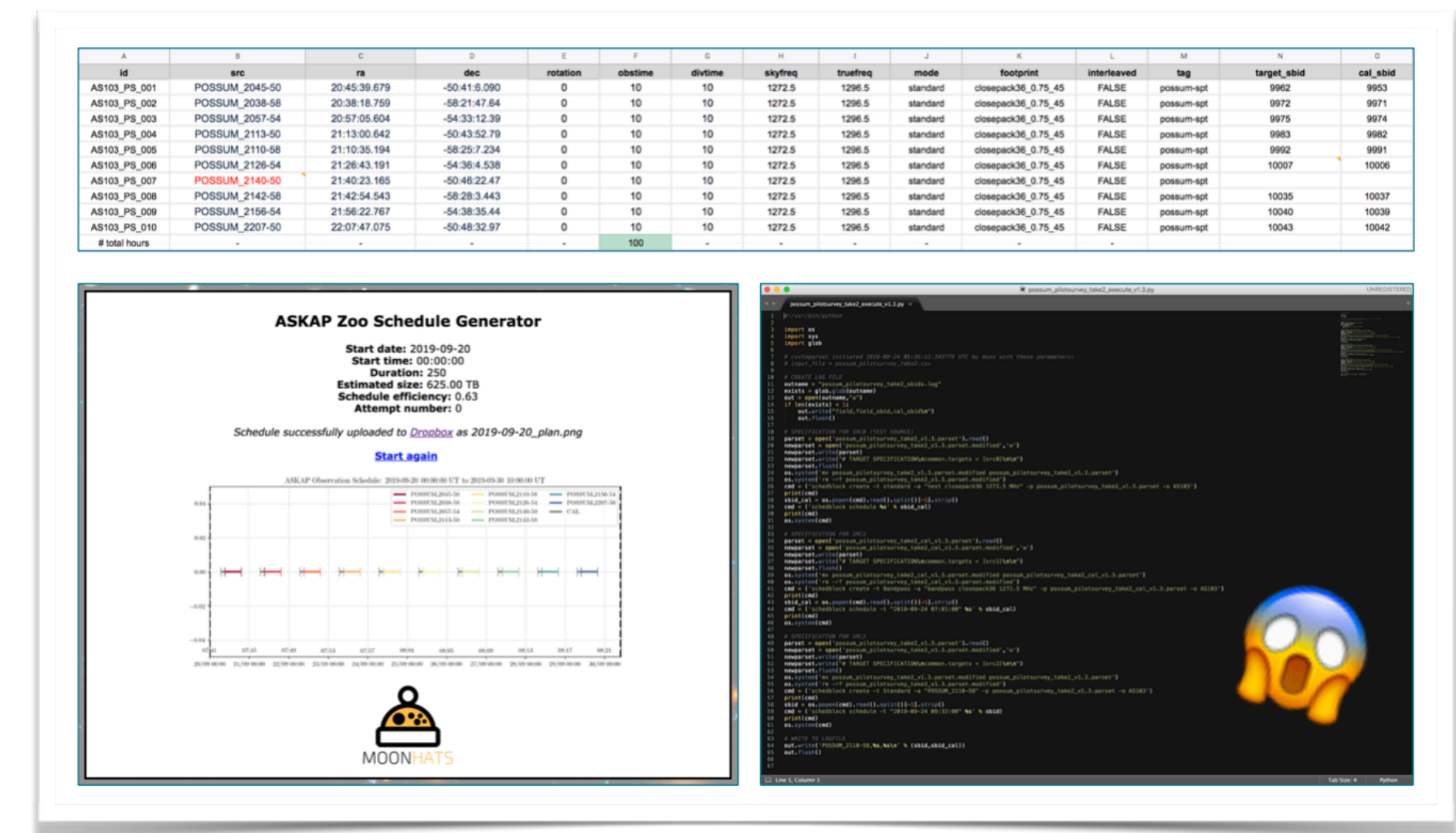
6. The shadowy place

The challenges we continue to face in operations



The shadowy place

- I've focused here on **key themes** that are important for operations, alongside **successes** - but there are significant **challenges** too!
- Automation** in the scope of a complex system: "hacky" code vs. the "proper solution"
- Distribution of **specialised expert knowledge** to non-experts for operations/diagnostics
- Keeping **communication channels** open
- Fractional** commitment and its role in the curse of the **95% finished** product or feature



```
AntennaWeightsUpdater
antenna completed duration error filename
ak01 true 66 186 PAF ports with gain change > 6.0 dB since ... SB08150_201903181045_1200M_36b_SB10098.ak01.wt...
ak02 141 PAF ports with gain change > 6.0 dB since ...
ak03 160 PAF ports with gain change > 6.0 dB since ...
ak04 187 PAF ports with gain change > 6.0 dB since ...
ak05 true 67 SB08150_201903181045_1200M_36b_SB10098.ak06.wt...
ak07 true 65 SB08150_201903181045_1200M_36b_SB10098.ak07.wt...
ak08 true 66 SB08150_201903181045_1200M_36b_SB10098.ak08.wt...
ak09 true 67 SB08150_201903181045_1200M_36b_SB10098.ak09.wt...
ak10 176 PAF ports with gain change > 6.0 dB since ...
ak11 188 PAF ports with gain change > 6.0 dB since ...
ak12 true 65 SB08150_201903181045_1200M_36b_SB10098.ak12.wt...
ak13 true 66 SB08150_201903181045_1200M_36b_SB10098.ak13.wt...
ak14 184 PAF ports with gain change > 6.0 dB since ...
ak15 true 64 SB08150_201903181045_1200M_36b_SB10098.ak15.wt...
ak16 true 65 SB08150_201903181045_1200M_36b_SB10098.ak16.wt...
ak17 179 PAF ports with gain change > 6.0 dB since ...
ak18 true 65 SB08150_201903181045_1200M_36b_SB10098.ak18.wt...
ak19 172 PAF ports with gain change > 6.0 dB since ...
ak20 187 PAF ports with gain change > 6.0 dB since ...
ak21 186 PAF ports with gain change > 6.0 dB since ...
ak22 10 PAF ports with gain change > 6.0 dB since o...
ak24 true 64 SB08150_201903181045_1200M_36b_SB10098.ak24.wt...
ak25 131 PAF ports with gain change > 6.0 dB since ...
ak26 true 65 SB08150_201903181045_1200M_36b_SB10098.ak26.wt...
ak27 true 63 SB08150_201903181045_1200M_36b_SB10098.ak27.wt...
ak28 187 PAF ports with gain change > 6.0 dB since ...
ak29 true 66 SB08150_201903181045_1200M_36b_SB10098.ak29.wt...
ak30 true 65 SB08150_201903181045_1200M_36b_SB10098.ak30.wt...
ak31 true 63 SB08150_201903181045_1200M_36b_SB10098.ak31.wt...
ak32 61 PAF ports with gain change > 6.0 dB since o...
ak33 188 PAF ports with gain change > 6.0 dB since ...
ak34 188 PAF ports with gain change > 6.0 dB since ...
ak35 188 PAF ports with gain change > 6.0 dB since ...
ak36 true 65 SB08150_201903181045_1200M_36b_SB10098.ak36.wt...

total_duration= 146
```


Conclusions

- Astronomy has entered a new era driven by **technology** and large amounts of **data**
- Radio observatory operations are in the **process** of adapting, but there is more to do!
- **Key:** people, automation, communication, overview, testing and the shadowy place
- We can **learn lots** from other observatories, other science fields and also industry
- Pathfinder telescope experience is **directly relevant** for next-generation facilities e.g. SKA

